Scalar-Flat Lorentzian Einstein-Weyl Spaces

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Abstract

We find all three-dimensional Einstein-Weyl spaces with the vanishing scalar curvature.

Three-dimensional Lorentzian Einstein metrics have constant curvature. In order to construct non-trivial gravitational models in three dimensions, one should therefore look at conformal geometries involving a non-metric connection. Einstein-Weyl geometries appear quite naturally in this context.

Let \mathcal{W} be a three-dimensional Weyl space i.e. a manifold with with a torsion-free connection D and a conformal metric [h] such that null geodesics of [h] are also geodesics for D. This condition is equivalent to $D_i h_{jk} = \omega_i h_{jk}$ for some one form ω . Here h_{jk} is a representative metric in the conformal class, and the indices i, j, k, ... go from 1 to 3. If we change this representative by $h \longrightarrow \phi^2 h$, then $\omega \longrightarrow \omega + 2d \ln \phi$. The one-form ω 'measures' the difference between D and the Levi-Civita connection ∇ of h. A tensor object T which transforms as $T \longrightarrow \phi^m T$ when $h_{ij} \longrightarrow \phi^2 h_{ij}$ is said to be conformally invariant of weight m. The formula for a covariant weighted derivative of a one-form of weight m is

$$D_i V_j = \nabla_i V_j + \frac{1}{2} ((1 - m)\omega_i V_j + \omega_j V_i - h_{ij}\omega_k V^k). \tag{1}$$

The Ricci tensor W_{ij} is related to the Ricci tensor R_{ij} and of ∇ by

$$W_{ij} = R_{ij} + \nabla_i \omega_j - \frac{1}{2} \nabla_j \omega_i + \frac{1}{4} \omega_i \omega_j + h_{ij} \left(-\frac{1}{4} \omega_k \omega^k + \frac{1}{2} \nabla_k \omega^k \right).$$

The conformally invariant Einstein–Weyl (EW) condition on (W, h, ω) is $W_{(ij)} = W h_{ij}/3$, or in terms of the Riemannian data:

$$\chi_{ij} := R_{ij} + \frac{1}{2} \nabla_{(i} \omega_{j)} + \frac{1}{4} \omega_{i} \omega_{j} - \frac{1}{3} \left(r + \frac{1}{2} \nabla^{k} \omega_{k} + \frac{1}{4} \omega^{k} \omega_{k} \right) h_{ij} = 0, \tag{2}$$

where χ_{ij} is the trace-free part of the Ricci tensor of the Weyl connection, and $r = h^{ij}R_{ij}$. Weyl spaces which satisfy (2) will be called Einstein-Weyl spaces. The EW equations can be regarded as an integrable system. This is because both the twistor theory [5] and the Lax representation [4] exist. One should therefore be able to construct large families of explicit solutions.

In this paper we shall find explicitly all EW spaces with vanishing scalar curvature $W = h^{ij}W_{ij}$. We shall first establish the following result:

Lemma 1 If the scalar curvature of the Weyl connection vanishes on and EW space (W, [h], D), then the Faraday two form $F_{ij} := \nabla_{[i}\omega_{j]}$ is null.

Proof. The Bianchi identities for the curvature of the Weyl connection written in terms of the Levi-Civita connection and ω are [6]

$$\nabla^i F_{ij} + \frac{1}{2} \omega^i F_{ij} + \frac{1}{3} (\nabla_j W + \omega_j W) = 0.$$
(3)

Assume W = 0 (this is a well defined condition as W is conformally invariant of weight -2). Contracting (3) with ∇^j , and using $\nabla^j \nabla^i F_{ij} = 0$ yields

$$0 = (\nabla^j \omega^i) F_{ij} + \omega^i \nabla^j F_{ij} = F^{ij} F_{ij} - \frac{1}{2} \omega^i \omega^j F_{ij},$$

so F is null.

We conclude that there are no non-trivial scalar-flat EW spaces in the Euclidean signature [1]. Non-trivial solutions can be found in the indefinite signature:

Proposition 2 Let (h, ω) be an Einstein-Weyl structure with vanishing scalar curvature. Then either (h, ω) is flat, or the signature is (++-) and there exist local coordinates $x^i = (y, x, t)$ such that $\omega = y dt$, and h is given by one of two solutions:

$$h_1 = dy^2 + 2dxdt + \left(x[R(t) - \frac{y}{2}] + \frac{1}{48}y^4 + \frac{1}{12}R(t)y^3 + S(t)y\right)dt^2,\tag{4}$$

$$h_2 = dy^2 + 2dxdt - \frac{4x}{y}dydt + \left(\frac{x^2}{y^2} + \frac{xy}{2} + \frac{1}{8}y^4 + R(t)y^2 + S(t)y\right)dt^2,$$
 (5)

where R(t) and S(t) are arbitrary functions with continuous second derivatives.

Proof. Lemma 1 implies that $F_{ij} = \nabla_{[i}\omega_{j]}$ is a closed null two-form. The conformal freedom together with the Darboux theorem imply the existence of coordinates such that $\omega_i = y\nabla_i t$. Therefore $\omega \wedge d\omega = 0$, and the nullity of F gives $*F \wedge \omega = 0$. We can rewrite Bianchi identity (3) as

$$2(d*F) + \omega \wedge *F = 0.$$

and deduce d*F = 0. Therefore $\nabla^i F_{ij} = 0$, and $\varepsilon_i^{jk} F_{jk} = f(t) \nabla_i t$. Redefining y, t we can set f(t) = 1. The most general metric consistent with $dt = *dy \wedge dt$ is

$$h = dy^2 + 2(\hat{E}ds + \hat{F}dy + \hat{G}dt)dt,$$

where \hat{E}, \hat{F} , and \hat{G} are functions of (s, y, t). Put $\hat{E} = \partial x/\partial s$ and define $G = (\hat{G} - x_t)/2, F = \hat{F} - x_y$, so that

$$h = dy^2 + 2dxdt + 2Fdydt + Gdt^2, \qquad \omega = ydt.$$
 (6)

The freedom $x \to x + P(y,t)$ implies that F(x,y,t), and G(x,y,t) are defined up to addition of derivatives of P(y,t). Furthermore the conformal scale is only fixed up to arbitrary functions of t, $h \mapsto \tilde{h} = \Omega h$. This leads to to the redefinitions $(x,y,t) \to (\tilde{x},\tilde{y},\tilde{t})$, given by

$$\tilde{t} = T(t), \qquad \tilde{y} = \frac{y}{T_t} - 2\frac{T_{tt}}{T_t}, \qquad \tilde{x} = \frac{x}{T_t^3} + P(y, t), \qquad \Omega = T_t^2.$$
 (7)

These transformations will latter be used to simplify F and G. Now impose the EW equations: Equation $\chi_{12} = 0$ implies $F_{xx} = 0$, and we can choose P(y,t) such that F = xl(y,t). Now $\chi_{12} = \chi_{22} = 0$

0. Equations $\chi_{11} = 0$, $\chi_{23} = 0$ are equivalent and imply $a_y = G_{xx}$. Take $G = x^2 l_y / 2 + x m(y, t) + n(y, t)$. The vanishing of the scalar curvature

$$W = r + 2\nabla^k \omega_k - \frac{1}{2}\omega^k \omega_k = (3l^2 - 6l_y)/2$$

gives $l^2 = 2l_y$, therefore l(y, t) = 0 (case **1**), or l(y, t) = -2/(y + c(t)) (case **2**).

- In case 1 $\chi_{31} = 0$ implies m(y,t) = R(t) y/2. Finally $\chi_{33} = 0$ yields $n(y,t) = y^4/48 + R(t)y^3/12 + S(t)y + Z(t)$). Redefining x, we set Z(t) = 0, and the metric is given by (4).
- In case 2 the conformal freedom (7) can be used to eliminate c(t). This is achieved by setting $T_{tt} = c(t)$, and redefining m(y,t), and n(y,t). Now $\chi_{31} = 0$ implies m(y,t) = y/2 + P(t), and $\chi_{33} = 0$ gives $n(y,t) = y^4/8 + y^3P(t)/4 + R(t)y^2 + S(t)y$. In fact we can get rid of P(t): Replace x by $x y^2w(t)/2$, and redefine R(t) to obtain (5).

The next proposition shows that there doesn't exist a combination of coordinate and conformal transformations which maps (h_1, ω) to (h_2, ω) . Cases **1** and **2** are essentially distinct and can be invariantly characterized:

Proposition 3 Let (h, ω) be a non-flat EW structure with a vanishing scalar curvature, and let F be a corresponding Faraday two form. If *F is parallel with respect to a weighted Weyl connection, then (h, ω) is locally given by (4). Otherwise it is locally given by (5).

Proof. In both cases 1 and 2 * $F = \varepsilon_i^{jk} F_{jk} dx^i = dt$. First notice that vanishing of D(fdt) for some f is invariant condition; in a conformal scale defined by f the one-form *F is parallel. Here we treat fdt as a weighted object. In case 1 we find that with m = 3/2 we have D(fdt) = 0 if f = f(t), and $f_t + fR/2 = 0$. In case 2 D(fdt) doesn't vanish for any f. For example $(D(fdt))_{13} = -f/y$.

Using the formula for a weighted derivative of a vector of weight m

$$D_i V^j = \nabla_i V^j - \frac{1}{2} \delta_i^j \omega_k V^k - \frac{m+1}{2} \omega_i V^j + \frac{1}{2} \omega^j V_i$$

we deduce that the EW structure (4) (case 1) admits a covariantly constant vector

$$V = \exp\left(-\frac{1}{2} \int_{-\infty}^{t} R(t) dt\right) \frac{\partial}{\partial x}$$

with weight -1/2. Therefore (h_1, ω) belongs to the conformal class of the dKP Einstein-Weyl spaces [4].

It is natural to ask if other special classes of Einstein–Weyl spaces include Lorentzian scalar-flat examples. Along these lines, we have the following observation.

Proposition 4 Let R be an arbitrary function of one variable. Then the Weyl structure

$$h = 4\frac{(z + R(v))^2}{(1 + vw)^2} dv dw + dz^2, \qquad \omega = \frac{4}{z + R(v)} dz$$
 (8)

is scalar-flat, and $u = 2\log(2(z+R(v))/(1+vw))$ is a solution of the Lorentzian $SU(\infty)$ Toda equation $4u_{vw} + (e^u)_{zz} = 0$.

Proof. These are straightforward verifications, using the fact that the scalar curvature of the Weyl structure $(h = e^u dv dw + dz^2, \omega = 2u_z dz)$ is $\frac{1}{2}u_{zz} + \frac{1}{4}u_z^2$: these spaces are Lorentzian analogues of the hyperCR–Toda Einstein–Weyl spaces [2], see also [7].

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